### **Joint Syllabis**

### "Problems of particle physics and cosmology"

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Degree program: 010975 "Physics of fundamental interactions" Specialization: 010975 "Physics of fundamental interactions" Chair "Fundamental interactions and Cosmology" 1th academic year of the Master of Science' degree, 2<sup>nd</sup> semester

> Head of Chair, academician RAS Victor A. Matveev

#### Lecture 1. Problems of the Hot Big Bang Theory (2 hours)

Horizon problem, entropy problem, flatness problem, origin of the primordial inhomogeneities.

#### Lecture 2. Inflation in slow roll regime (4 hours)

Inflation as a solution to the Hot Big Bang problems; the slow roll conditions; chaotic, new, hybrid inflationary scenarios.

#### Lecture 3. Gaussian random variables and Gaussian random fields (2 hours)

Propeties of Gaussian random variables, Gaussian random fields.

#### Lecture 4. Generation of cosmological perturbations at inflation (4 hours)

Inflaton field fluctuations; generation of primordial scalar perturbations, generation of primordial tensor perturbations (gravitational waves); amplitudes and tilts of the perturbation power spectra.

#### Lecture 5. Particle production in background fields (2 hours)

Method of the Bogoliubov transformation: bosons, fermions.

#### Lecture 6. Postinflationary reheating (4 hours)

Perturbative mechanism: decay of the inflaton field oscillations; thermalisation of particles in the expanding Universe; parametric resonance in the inflaton decay: decay of large amplitude oscillations, production of heavy fermions.

#### Lecture 7. Jeans instability (2 hours)

Jeans instability in static background, development of the instability in the expanding Universe, large scale structures in the Universe.

## Lecture 8. Cosmological perturbations in linear approximation (4 hours)

Linearized energy-momentum tensor of ideal fluid; linearized Einstein equations; helicity decomposition: scalar, vector, tensor modes.

## Lecture 9. Evolutions of vector and tensor modes (2 hours)

Superhorizon and subhorizon modes; matching at horizon.

## Lecture 10. Scalar perturbation modes of single-component fluids (2 hours)

Radiation domination; matter domination; matter perturbation at dark energy dominated stage.

## Lecture 11. Large scale structure formation in the Universe (2 hours)

Linear stage of scalar perturbation evolution after recombination; power spectrum; entrance to nonlinear regime; mass spectrum of the large scale structures.

## Lecture 12. Cosmic microwave background anisotropy (2 hours)

Anisotropy of CMB within the photon instant decoupling approximation.

## References

### **Basic references**

- **1.** Introduction to the theory of the early Universe: Cosmological perturbations and inflationary theory. Dmitry S. Gorbunov, Valery A. Rubakov, World Scientific (2011) 489 p
- 2. Introduction to the theory of the early Universe: Hot big bang theory. Dmitry S. Gorbunov, Valery A. Rubakov, World Scientific (2011) 473 p
- 3. Modern Cosmology, Scott Dodelson, Amsterdam: Academic Press (2003) 440 p
- **4.** Physical Foundations of Cosmology», Viatcheslav Mukhanov, Cambridge University Press (2005) 421 p

# **Additional references**

- 1. Cosmology, Steven Weinberg, Oxford University Press (2008) 593 p
- 2. Particle Physics and Inflationary Cosmology, Contemp.Concepts Phys. 5 (1990) 1-362
- **3.** The Early Universe, Edward W. Kolb, Michael S. Turner, Redwood City: Addison-Wesley (1990) 547 p